Partial differential equations

# Stability for entire radial solutions to the biharmonic equation with negative exponents 

# Stabilité des solutions radiales entières de l'équation biharmonique avec exposants négatifs 

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## A B S T R A C T

In this note, we are interested in entire solutions to the semilinear biharmonic equation

$$
\Delta^{2} u=-u^{-p}, u>0 \quad \text { in } \mathbb{R}^{N}
$$

where $p>0$ and $N \geq 3$. In particular, the stability outside a compact set of the entire radial solutions will be completely studied, which resolves the remaining case in [5].
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## R É S U M É

Dans cette note, on s'intéresse aux solutions radiales entières de l'équation semilinéaire biharmonique

$$
\Delta^{2} u=-u^{-p}, \quad u>0 \quad \text { dans } \mathbb{R}^{N}
$$

où $p>0$ et $N \geq 3$. En particulier, on étudie la stabilité en dehors d'un compact des solutions radiales entières, et on résout un cas ouvert dans [5].
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## 1. Introduction

In this note, we are interested in entire radial solutions to the biharmonic equation

$$
\begin{equation*}
\Delta^{2} u=-u^{-p}, u>0 \quad \text { in } \mathbb{R}^{N} \tag{1.1}
\end{equation*}
$$

where $p>0$ and $N \geq 3$.

[^0]Recently, the fourth-order equations have attracted the interest of many researchers. In particular, the existence, multiplicity, stability, and qualitative properties of solutions to equation (1.1) are studied in many works, especially for radial solutions. It has been proved in [6] that, if $0<p \leq 1$, the equation (1.1) admits no entire smooth solution. It is showed in [4,7] that, for any $p>1$, there exist radial solutions to (1.1).

Definition 1. A solution $u$ to (1.1) is said stable in $\Omega \subseteq \mathbb{R}^{N}$ if there holds

$$
\int_{\Omega}|\Delta \phi|^{2} \mathrm{~d} x-p \int_{\Omega} u^{-p-1} \phi^{2} \mathrm{~d} x \geq 0 \quad \text { for any } \phi \in C_{0}^{\infty}(\Omega)
$$

Moreover, a solution $u$ to (1.1) is said stable outside a compact set $K$ if $u$ is stable in $\mathbb{R}^{N} \backslash K$. For simplicity, we say also that $u$ is stable if $\Omega=\mathbb{R}^{N}$.

We consider the following initial value problem

$$
\left\{\begin{array}{l}
\Delta^{2} u=-u^{-p} \quad \text { for } r \in\left[0, R_{\alpha, \beta}\right)  \tag{1.2}\\
u^{\prime}(0)=u^{\prime \prime \prime}(0)=0 \\
u(0)=\alpha, \quad \Delta u(0)=\beta
\end{array}\right.
$$

for any $\alpha, \beta \in \mathbb{R}$, we denote by $u_{\alpha, \beta}$ the (local) solution to (1.2) and by [ $0, R_{\alpha, \beta}$ ) the maximal interval of existence. Notice that the equation (1.2) is invariant under the scaling transformation

$$
u_{\lambda}(x)=\lambda^{-\frac{4}{p+1}} u(\lambda x), \lambda>0
$$

Therefore, we need only to consider the case $\alpha=1$. We will denote $u_{1, \beta}$ by $u_{\beta}$. Let $p>1$, it is known from [3,5,7] that

- there is no global solution to (1.2) if $N \leq 2$;
- for $N \geq 3$, there exists $\beta_{0}>0$ depending on $N$ such that the solution to (1.2) is globally defined if and only if $\beta \geq \beta_{0}$. Furthermore, $\lim _{r \rightarrow \infty} \Delta u_{\beta} \geq 0$ and $\lim _{r \rightarrow \infty} \Delta u_{\beta}=0$ if and only if $\beta=\beta_{0}$;
- for $N \geq 3$, any entire solution $u_{\beta}$ is stable outside a compact set if $\beta>\beta_{0}$;
- for $N=4, u_{\beta_{0}}$ is unstable outside every compact set;
- for $5 \leq N \leq 12$, there exists a critical value $p_{N}>1$ (see below for the precise definition) such that, if $1<p \leq p_{N}, u_{\beta}$ is stable for every $\beta \geq \beta_{0}$, while for $p>p_{N}$, there exists $\beta_{1}>\beta_{0}$ such that $u_{\beta}$ is stable if and only if $\beta \geq \beta_{1}$, and $u_{\beta_{0}}$ is unstable outside every compact set;
- for $N \geq 13$ and any $p>1, u_{\beta}$ is stable for every $\beta \geq \beta_{0}$.

Moreover, Warnault [8] proved that equation (1.1) admits no stable solution (radial no not) for $N \leq 4$. So it remains to consider the eventual stability outside a compact set for $N=3$ and $\beta=\beta_{0}$.

The stability property of entire radial solutions is closely related to their asymptotic behaviors. Let us recall the asymptotic behaviors showed in $[2,3,5]$. For $N=3$ and $\beta=\beta_{0}$, the following hold:

$$
\begin{cases}\lim _{r \rightarrow \infty} u_{\beta_{0}}(r) r^{-1}=\ell>0, & \text { if } p>3  \tag{1.3}\\ \lim _{r \rightarrow \infty} u_{\beta_{0}}(r) r^{-1}(\ln r)^{-\frac{1}{4}}=\sqrt[4]{2}, & \text { if } p=3 \\ \lim _{r \rightarrow \infty} u_{\beta_{0}}(r) r^{-\frac{4}{p+1}}=\left[-Q_{4}\left(-\frac{4}{p+1}\right)\right]^{-\frac{1}{p+1}}=: L_{0}, & \text { if } 1<p<3\end{cases}
$$

where $Q_{4}$ is defined by

$$
\begin{equation*}
Q_{4}(m):=m(m+2)(N-2-m)(N-4-m) \tag{1.4}
\end{equation*}
$$

Remark that equation (1.1) has a singular solution $u_{s}(r) \equiv L_{0} r^{\frac{4}{p+1}}$, if $Q_{4}\left(-\frac{4}{p+1}\right)<0$.
From [2], we know that for $N=3$, there exist $3>p_{c}^{+}>p_{c}>1$ such that, if $p=p_{c}$ or $p=p_{c}^{+}$, then $-p Q_{4}(m)=\frac{9}{16}$ with $m=-\frac{4}{p+1}$, and if $p_{c}<p<p_{c}^{+}$then $-p Q_{4}(m)>\frac{9}{16}$. For $N \geq 5, p_{N}$ is the unique root of

$$
-p Q_{4}\left(-\frac{4}{p+1}\right)=\frac{N^{2}(N-4)^{2}}{16}
$$

in $(1, \infty)$.

Theorem A (Theorem 1.6 in [5]). Let $N=3, p>1$. We have:
(i) if $p_{c}^{+}<p<3$ or $1<p<p_{c}$, then $u_{\beta_{0}}$ is stable outside a compact set;
(ii) if $p_{c}<p<p_{c}^{+}, u_{\beta_{0}}$ is unstable outside every compact set;
(iii) if $p \geq 3$, then $u_{\beta_{0}}$ is stable outside a compact set.

Open problem: What is the stability behavior outside compact set when $\beta=\beta_{0}, N=3, p=p_{c}$ or $p=p_{c}^{+}$?
The following result gives the definite answer.

Theorem 1.1. Let $N=3, p=p_{c}$ or $p=p_{c}^{+}$, the solution $u_{\beta_{0}}$ to equation (1.2) is stable outside a compact set.

Indeed, we will prove a refined asymptotic behavior for the radial solution $u_{\beta_{0}}$ and use the following Hardy-Rellich inequality with weights, see Corollary 5.4 in [1].

Lemma 1.2. Let $N \geq 3, \Omega=\mathbb{R}^{N} \backslash B_{1}$, then the following inequality holds

$$
\begin{align*}
\int_{\Omega}|\Delta \phi|^{2} \mathrm{~d} x & -\frac{N^{2}(N-4)^{2}}{16} \int_{\Omega} \frac{\phi^{2}}{|x|^{4}} \mathrm{~d} x \\
& \geq \frac{N^{2}-4 N+8}{8} \int_{\Omega} \frac{\phi^{2}}{|x|^{4} \ln ^{2}|x|} \mathrm{d} x+\frac{9}{16} \int_{\Omega} \frac{\phi^{2}}{|x|^{4} \ln ^{4}|x|} \mathrm{d} x, \quad \forall \phi \in C_{c}^{\infty}(\Omega) . \tag{1.5}
\end{align*}
$$

## 2. Proof of Theorem 1.1

Rewrite the equation (1.1) with the radial coordinate.

$$
u^{(4)}+\frac{2(N-1)}{r} u^{\prime \prime \prime}+\frac{(N-1)(N-3)}{r^{2}} u^{\prime \prime}-\frac{(N-1)(N-3)}{r^{3}} u^{\prime}=-u^{-p}
$$

Denote $\alpha:=-m=\frac{4}{p+1}$. Without confusion, from now on we omit the index $\beta_{0}$ and fix $N=3, p \in(1,3)$. Let $v(t)=$ $r^{-\alpha} u-L_{0}$ with $t=\ln r$, then $v$ satisfies

$$
\begin{equation*}
v^{(4)}+2(2 \alpha-1) v^{\prime \prime \prime}+\left(6 \alpha^{2}-6 \alpha-1\right) v^{\prime \prime}+2(2 \alpha-1)\left(\alpha^{2}-\alpha-1\right) v^{\prime}-(p+1) L_{0}^{-(p+1)} v+g(v)=0 \tag{2.1}
\end{equation*}
$$

where $g(v)=\left(v+L_{0}\right)^{-p}-L_{0}^{-p}+p L_{0}^{-(p+1)} v$. As $1<p<3$, by (1.3), we have $\lim _{t \rightarrow \infty} v(t)=0$, so $g(v)=O\left(v^{2}\right)$ as $t \rightarrow \infty$.
The corresponding characteristic polynomial of equation (2.1) is

$$
\lambda^{4}+2(2 \alpha-1) \lambda^{3}+\left(6 \alpha^{2}-6 \alpha-1\right) \lambda^{2}+2(2 \alpha-1)\left(\alpha^{2}-\alpha-1\right) \lambda-(p+1) L_{0}^{-(p+1)}=0
$$

Using MATLAB, we have the following four roots of the above polynomial:

$$
\left\{\begin{array}{l}
\lambda_{1}=\frac{1}{2}-\alpha+\frac{1}{2} \sqrt{5+4 \sqrt{h(p, \alpha)}} \\
\lambda_{2}=\frac{1}{2}-\alpha-\frac{1}{2} \sqrt{5+4 \sqrt{h(p, \alpha)}} \\
\lambda_{3}=\frac{1}{2}-\alpha+\frac{1}{2} \sqrt{5-4 \sqrt{h(p, \alpha)}} \\
\lambda_{4}=\frac{1}{2}-\alpha-\frac{1}{2} \sqrt{5-4 \sqrt{h(p, \alpha)}}
\end{array}\right.
$$

where $h(p, \alpha)=1+p \alpha(2-\alpha)(1+\alpha)(\alpha-1)$.
Recall that for $p=p_{c}$ or $p_{c}^{+}$, there holds $-p Q_{4}(-\alpha)=\frac{9}{16}$, i.e. $p \alpha(2-\alpha)(\alpha+1)(\alpha-1)=\frac{9}{16}$. Hence $h(p, \alpha)=\frac{25}{16}$ and

$$
\lambda_{1}=\frac{1}{2}-\alpha+\frac{1}{2} \sqrt{10}, \quad \lambda_{2}=\frac{1}{2}-\alpha-\frac{1}{2} \sqrt{10}, \quad \lambda_{3}=\lambda_{4}=\frac{1-2 \alpha}{2}, \quad \text { if } p=p_{c} \text { or } p_{c}^{+}
$$

As $\alpha \in(1,2)$ for $1<p<3$, we see that $\lambda_{1}>0, \lambda_{2}<\lambda_{3}=\lambda_{4}<0$. By the variation of parameters method, the solution $v$ to (2.1) is given by

$$
\begin{aligned}
v(t)= & \sum_{i=1}^{3} A_{i} \mathrm{e}^{\lambda_{i} t}+\sum_{i=1}^{3} B_{i} \int_{0}^{t} \mathrm{e}^{\lambda_{i}(t-s)} g(v(s)) \mathrm{d} s+A_{4} t \mathrm{e}^{\lambda_{4} t}+B_{4} \int_{0}^{t}(t-s) \mathrm{e}^{\lambda_{4}(t-s)} g(v(s)) \mathrm{d} s \\
= & A_{1}^{\prime} \mathrm{e}^{\lambda_{1} t}+A_{2} \mathrm{e}^{\lambda_{2} t}+A_{3} \mathrm{e}^{\lambda_{3} t}+A_{4} t \mathrm{e}^{\lambda_{4} t}-B_{1} \int_{t}^{\infty} \mathrm{e}^{\lambda_{1}(t-s)} g(v(s)) \mathrm{d} s \\
& +\sum_{i=2}^{3} B_{i} \int_{0}^{t} \mathrm{e}^{\lambda_{i}(t-s)} g(v(s)) \mathrm{d} s+B_{4} \int_{0}^{t}(t-s) \mathrm{e}^{\lambda_{4}(t-s)} g(v(s)) \mathrm{d} s
\end{aligned}
$$

where we used the fact that $\mathrm{e}^{-\lambda_{1} s} g(v(s)) \in L^{1}\left(\mathbb{R}_{+}\right)$. As $\lim _{t \rightarrow \infty} v(t)=0$ and $\lambda_{1}>0$, there holds $A_{1}^{\prime}=0$. Therefore, for any $\epsilon \in\left(0,-\lambda_{4}\right)$, there exists $C_{\epsilon}>0$ such that, for all $t \geq 0$,

$$
|v(t)| \leq C_{\epsilon} \mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}+C_{\epsilon} \int_{t}^{\infty} \mathrm{e}^{\lambda_{1}(t-s)}|g(v)(s)| \mathrm{d} s+C_{\epsilon} \int_{0}^{t} \mathrm{e}^{\left(\lambda_{4}+\epsilon\right)(t-s)}|g(v)(s)| \mathrm{d} s .
$$

Moreover, for any $\delta>0$, there exists $M>0$ such that $|g(v)(s)| \leq \delta|v(s)|$ if $s \geq M$. Then for $t \geq M$,

$$
\begin{align*}
|v(t)| & \leq O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)+C_{\epsilon} \delta \int_{t}^{\infty} \mathrm{e}^{\lambda_{1}(t-s)}|v(s)| \mathrm{d} s+C_{\epsilon} \delta \int_{M}^{t} \mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}|v(s)| \mathrm{d} s,  \tag{2.2}\\
& =O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)+C_{\epsilon} \delta K_{1}(t)+C_{\epsilon} \delta K_{2}(t)
\end{align*}
$$

with

$$
K_{1}(t):=\int_{M}^{t} \mathrm{e}^{\left(\lambda_{4}+\epsilon\right)(t-s)}|v(s)| \mathrm{d} s, \quad K_{2}(t):=\int_{t}^{\infty} \mathrm{e}^{\lambda_{1}(t-s)}|v(s)| \mathrm{d} s
$$

Thanks to (2.2), if we fix $\delta>0$ small enough such that $2 C_{\epsilon} \delta \leq \min \left(\lambda_{1},-\lambda_{4}-\epsilon\right)$, there holds

$$
\begin{aligned}
\left(K_{1}-K_{2}\right)^{\prime}(t) & =2|v(t)|+\left(\lambda_{4}+\epsilon\right) K_{1}(t)-\lambda_{1} K_{2}(t) \\
& \leq 2 C_{\epsilon} \delta\left(K_{1}+K_{2}\right)+\left(\lambda_{4}+\epsilon\right) K_{1}(t)-\lambda_{1} K_{2}+O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right) \\
& \leq O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)
\end{aligned}
$$

Using again $\lim _{t \rightarrow \infty} v(t)=0$, we have readily

$$
\lim _{t \rightarrow \infty} K_{1}(t)=\lim _{t \rightarrow \infty} K_{2}(t)=0
$$

Hence, $\left(K_{2}-K_{1}\right)(t) \leq O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)$ as $t \rightarrow \infty$. Going back to (2.2),

$$
\begin{equation*}
|v(t)| \leq O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)+2 C_{\epsilon} \delta K_{1}(t) \tag{2.3}
\end{equation*}
$$

Consequently,

$$
K_{1}^{\prime}(t)=|v(t)|+\left(\lambda_{4}+\epsilon\right) K_{1}(t) \leq O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon\right) t}\right)+\left(2 C_{\epsilon} \delta+\lambda_{4}+\epsilon\right) K_{1}(t)
$$

So $K_{1}(t)=O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon+2 C_{\epsilon} \delta\right) t}\right)$; we get $|v(t)|=O\left(\mathrm{e}^{\left(\lambda_{4}+\epsilon+2 C_{\epsilon} \delta\right) t}\right)$ by (2.3). Let $\sigma=-\lambda_{4}-\epsilon-2 C_{\epsilon} \delta>0$, we obtain

$$
\begin{equation*}
u(r)=L_{0} r^{\alpha}+r^{\alpha} O\left(r^{-\sigma}\right), \quad \text { as } r \rightarrow \infty \tag{2.4}
\end{equation*}
$$

Finally, let $R>0$ be large enough, we apply Lemma 1.2 with $N=3$. Recall that $p=p_{c}$ or $p_{c}^{+}$, for any $\phi \in C_{c}^{\infty}\left(\mathbb{R}^{3} \backslash B_{R}\right)$, we have then

$$
\begin{aligned}
& \int_{\mathbb{R}^{3} \backslash B_{R}}|\Delta \phi|^{2} \mathrm{~d} x-p \int_{\mathbb{R}^{3} \backslash B_{R}} u^{-p-1} \phi^{2} \mathrm{~d} x \\
\geq & \int_{\mathbb{R}^{3} \backslash B_{R}}|\Delta \phi|^{2} \mathrm{~d} x-p \int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4}\left[L_{0}^{-(p+1)}-O\left(r^{-\sigma}\right)\right] \phi^{2} \mathrm{~d} x \\
= & \int_{\mathbb{R}^{3} \backslash B_{R}}|\Delta \phi|^{2} \mathrm{~d} x-p L_{0}^{-(p+1)} \int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4} \phi^{2}-\int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4} O\left(r^{-\sigma}\right) \phi^{2} \mathrm{~d} x \\
= & \int_{\mathbb{R}^{3} \backslash B_{R}}|\Delta \phi|^{2} \mathrm{~d} x+p Q_{4}\left(-\frac{4}{p+1}\right) \int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4} \phi^{2}-\int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4} O\left(r^{-\sigma}\right) \phi^{2} \mathrm{~d} x \\
= & \int_{\mathbb{R}^{3} \backslash B_{R}}|\Delta \phi|^{2} \mathrm{~d} x-\frac{9}{16} \int_{\mathbb{R}^{3} \backslash B_{R}} r^{-4} \phi^{2}-\int_{\mathbb{R}^{3} \backslash B_{R}} O\left(r^{-4-\sigma}\right) \phi^{2} \mathrm{~d} x \geq 0 .
\end{aligned}
$$

This implies that $u$ is stable outside a compact set. The proof is completed.

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## References

[1] P. Caldiroli, R. Musina, Rellich inequalities with weights, Calc. Var. Partial Differ. Equ. 45 (1-2) (2012) 147-164.
[2] J. Dávila, I. Flores, I. Guerra, Multiplicity of solutions for a fourth order equation with power-type nonlinearity, Math. Ann. 348 (1) (2010) 143-193.
[3] I. Guerra, A note on nonlinear biharmonic equations with negative exponents, J. Differ. Equ. 253 (11) (2012) 3147-3157.
[4] T. Kusano, M. Naita, C.A. Swanson, Radial entire solutions of even order semilinear elliptic equations, Can. J. Math. 40 (1988) 1281-1300.
[5] B.S. Lai, The regularity and stability of solutions to semilinear fourth-order elliptic problems with negative exponents, Proc. R. Soc. Edinb., Sect. A 146 (1) (2016) 195-212.
[6] B.S. Lai, D. Ye, Remarks on entire solutions for two fourth-order elliptic problems, Proc. Edinb. Math. Soc. 59 (3) (2016) 777-786.
[7] P.J. McKenna, W. Reichel, Radial solutions of singular nonlinear biharmonic equations and applications to conformal geometry, Electron. J. Differ. Equ. 2003 (37) (2003), 13 p.
[8] G. Warnault, Liouville theorems for stable radial solutions for the biharmonic operator, Asymptot. Anal. 69 (2010) 87-98.


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