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Small vibrations of a linearly elastic body surrounded by heavy, incompressible, non-Newtonian fluids with free surfaces

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Abstract

We consider the small transient motions of a coupled system constituted by a linearly elastic body and two heavy, incompressible, non-Newtonian fluids. Through a formulation in terms of non-linear evolution equations in Hilbert spaces of possible states with finite mechanical energy, we obtain existence and uniqueness results and study the influence of gravity. **To cite this article:** *C. Licht, Tran Thu Ha, C. R. Mécanique 333 (2005).*

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Résumé

Petites vibrations d'un corps linéairement élastique baigné par des fluides non-Newtoniens, incompressibles, pesants, à surfaces libres. On considère les petits mouvements instationnaires d'un système couplé constitué d'un solide linéairement élastique et de deux fluides non-Newtoniens, incompressibles, pesants à surfaces libres. Une formulation en termes d'équations d'évolution non-linéaires dans des espaces de Hilbert d'états possibles d'énergie mécanique finie permet d'obtenir des résultats d'existence et d'unicité et d'étudier l'influence de la gravité. **Pour citer cet article :** *C. Licht, Tran Thu Ha, C. R. Mécanique 333 (2005).*

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1. Introduction

Some problems of offshore engineering lead us to consider the *small* transient vibrations of a coupled fluids–structure system around an equilibrium configuration which can be described as follows. A linearly elastic body occupies the closure \bar{B} of a domain B of \mathbb{R}^3 with a Lipschitz-continuous boundary ∂B . The heavy, incompressible and not necessarily Newtonian viscous fluids fill up the reunion Ω of two disjoint connected open sets of \mathbb{R}^3 with Lipschitz boundaries Ω^i (the inner fluid) and Ω^e (the outer fluid); Ω^i is bounded and Ω^e may be unbounded, they respectively lie in the half space $\{x_3 \leq h^i\}$ and in the strip $\{d^e \leq x_3 \leq h^e\}$. Parts of the boundaries $\partial\Omega^i$ and $\partial\Omega^e$ of Ω^i and Ω^e , the so-called free surfaces S_F^i and S_F^e , are included in the planes $\{x_3 = h^i\}$ and $\{x_3 = h^e\}$, let $S_F = S_F^i \cup S_F^e$. The body domain and the fluid region share a boundary S_w , the wet surfaces, let $S_w^i = S_w \cap \partial\Omega^i$ and $S_w^e = S_w \cap \partial\Omega^e$. Whereas the fluid particles adhere perfectly to a rigid ‘bottom’ $S_B = \partial\Omega^e \setminus (S_w^e \cup S_F^e)$, we have $\partial\Omega^i = S_w^i \cup S_F^i$. The body is respectively clamped and free of surface loading on the remaining parts Γ_0 and Γ_1 of ∂B . The state of the coupled system is determined by the quadruplet $u_g = (v_f, \eta, s, v_b)$, v_f is the fluid velocity, η the free surface elevation, s and v_b the fields of displacement and velocity in the body; we will use the index g to specify that gravity effects are taken into account (see Section 3.2 also). Let σ_f and σ_b the (fluctuations of) stresses in the fluid and in the body, the equations of the small motions may be expressed as:

$$\rho_f \frac{\partial v_f}{\partial t} - \operatorname{div} \sigma_f = f_f \quad \text{in } \Omega \quad (1)$$

$$\sigma_f n = -\rho_f g \eta n \quad \text{on } S_F \quad (2)$$

$$\frac{\partial \eta}{\partial t} = v_f \cdot n \quad \text{on } S_F \quad (3)$$

$$v_f = 0 \quad \text{on } S_B \quad (4)$$

$$\sigma_f n = \sigma_b n \quad \text{on } S_w \quad (5)$$

$$v_f = v_b \quad \text{on } S_w \quad (6)$$

$$\rho_b \frac{\partial v_b}{\partial t} - \operatorname{div} \sigma_b = f_b \quad \text{in } B \quad (7)$$

$$v_b = \frac{\partial s}{\partial t} \quad \text{in } B \quad (8)$$

$$s = 0 \quad \text{on } \Gamma_0 \quad (9)$$

$$\sigma_b v = 0 \quad \text{on } \Gamma_1 \quad (10)$$

Here, ρ_f , ρ_b are the densities in the fluids and in the body, g is the gravity acceleration, f_f , f_b are the densities of the applied body forces excluding the own-weights, n , v are the unit normals outward Ω and B , and t denotes the time. The density is constant in each fluid while ρ_b and $1/\rho_b$ are positive elements of $L^\infty(B)$. The constitutive equations are:

$$\operatorname{div} v_f = 0, \quad (\sigma_f)_{\operatorname{dev}} \in \partial d(x, e(v_f)) \quad \text{in } \Omega \quad (11)$$

$$\sigma_b = a e(s) \quad \text{in } B \quad (12)$$

where x is the space variable, e stands for the symmetrized gradient, $(\sigma_f)_{\operatorname{dev}}$ is the deviatoric part of σ_f , a is the linearized tensor of elasticity and ∂d denotes the sub-differential of the convex density function of dissipation potential $d(x, \cdot)$. We assume that a enjoys the usual uniform properties of symmetry, boundedness and ellipticity, while $d(\cdot, e)$ is constant in each fluid and $d(x, \cdot)$ is a convex lower semi-continuous function in the set S_{dev}^3 of 3×3 symmetric matrices with vanishing traces, such that:

$$d(x, 0) = 0, \quad \forall x \in \Omega \quad \exists c > 0, \quad p > 1; \quad d(x, q) \geq c|q|^p, \quad \forall q \in S_{\operatorname{dev}}^3, \quad \forall x \in \Omega \quad (13)$$

An initial state u_g^0 being given, the problem is to find $u_g(t)$, $t \in [0, T]$, satisfying (1)–(12). In [1,2], similar simplified situations were considered with Newtonian fluids, here the dissipation potential is not necessarily quadratic.

As previously, it can be shown that the problem is formally equivalent to

$$\frac{du_0}{dt} + A_0 u_0 \ni (f_f^*/\rho_f, 0, f_b/\rho_b), \quad u_0(0) = u_0^0 \tag{25}$$

and that A_0 is maximal-monotone, thus (25) admits a unique solution enjoying regularity properties analogous to those of u_g in (20). Under the additional assumption on d :

$$\exists C > 0; \quad d(x, q) \leq C(1 + |q|)^p, \quad \forall q \in S_{\text{dev}}^3, \quad \forall x \in \Omega \tag{26}$$

we can precise the status of this simplified modelling with respect to the modelling of Section 2. Introducing a family of linear isometric operators P_g :

$$H_0 \ni u_0 = (v_f, s, v_b) \longmapsto P_g u_0 = (v_f, 0, s, v_b) \in H_g \tag{27}$$

we have the following approximation result:

Theorem 3.1. *If $\lim_{g \rightarrow 0} |P_g u_0^0 - u_g^0|_g = 0$, then, when g goes to 0, the unique solution u_g of (20) converges to the unique solution u_0 of (25), in the sense that $\lim_{g \rightarrow 0} |P_g u_0(t) - u_g(t)|_g = 0$ uniformly on $[0, T]$.*

From an easy non-linear extension of the Trotter theory [5] of approximation of semi-groups of operators acting on variable Hilbert spaces, the proof of Theorem 3.1 reduces to the proof of

$$\lim_{g \rightarrow 0} |P_g (I + \lambda A_0)^{-1} f - (I + \lambda A_g)^{-1} P_g f|_g = 0, \quad \forall (\lambda, f) \in (0, \infty) \times H_0$$

which is a straightforward consequence of the obvious Γ -convergence of J_g to J_0 with respect to the sequential weak convergence on W_p .

Of course, the gravity acceleration does not go to zero, but a preliminary non-dimensional setting of the problem will involve a coefficient which increases from 0 with g . Thus, the practical interest of the previous theorem is to describe what is happening when this coefficient goes to zero: convergences in energy norms of the displacement and of the velocities, convergence to zero of the gravity potential energy of the fluid.

It is interesting to note that the problem at $g = 0$ may also concern the small motions of a deformable composite whose constituents, perfectly stuck together, are either linearly elastic or viscoplastic.

3.3. Taking into account fluctuations of atmospheric pressure and surface forces on the body

This realistic situation can be handled also by the tool of evolution equations governed by time-independent maximal-monotone operators in the case of Newtonian fluids with, for instance, dissipation density functions like $d(x, e) = \mu(x)|e|^2$, where μ is constant and positive in each fluid. More precisely, Eqs. (2) and (10) are replaced by:

$$\sigma_f n = -\rho_f g \eta n + \gamma \quad \text{on } S_F \tag{2'}$$

$$\sigma_b v = \delta \quad \text{on } \Gamma_1 \tag{10'}$$

If we assume that (γ, δ) belongs to $W^{2,\infty}(0, T; L^2(S_F)^3 \times L^2(\Gamma_1)^3)$, then there exists a unique (ω_f, ω_b) in $W^{2,\infty}(0, T; V_2)$ such that:

$$\int_{\Omega} 2\mu e(\omega_f) \cdot e(\varphi) \, dx + \int_B a e(\omega_b) \cdot e(\psi) \, dx = \int_{S_F} \gamma \cdot \varphi \, ds + \int_{\Gamma_1} \delta \cdot \psi \, ds \quad \forall (\varphi, \psi) \in V_2 \tag{28}$$

In this way, $\bar{u}_g := u_g - (\omega_f, 0, \omega_b, \omega_b)$ satisfies an evolution equation like (20), but with $(f_f^*/\rho_f - d\omega_f/dt, \omega_f \cdot n, \omega_b - d\omega_b/dt, f_b/\rho_b - d\omega_b/dt)$ as second member and $\bar{u}_g^0 := u_g^0 - (\omega_f(0), 0, \omega_b(0), \omega_b(0))$ as initial data, so that the additional assumptions $\bar{u}_g^0 \in \text{dom}(A_g)$ and $(f_f, f_b) \in W^{1,\infty}(0, T; L^2(\Omega)^3 \times L^2(B)^3)$ imply existence and uniqueness of u_g in $C^1([0, T]; H_g) \cap C^0([0, T]; \text{dom}(A_g))$.

3.4. Superposed given rigid motion

We assume that the given motion of the frame attached to the body is such that the concept of mean reference configuration is meaningful. In that case, the small motions around this configuration may be described by:

$$\frac{du_g}{dt} + A_g u_g \ni (f_f^*/\rho_f, 0, 0, f_b/\rho_b) + L(t)u_g + c(t), \quad u_g(0) = u_g^0 \quad (29)$$

where the element c of H_g and the element L of the space $\mathcal{L}(H_g)$ of bounded linear operators on H_g account for the additional acceleration terms in (1), (7). If we assume that the given motion of the frame is of class $W^{2,\infty}(0, T)$ then $L \in W^{2,\infty}(0, T; \mathcal{L}(H_g))$ and $c \in W^{1,\infty}(0, T; H_g)$, thus existence and uniqueness of a solution of (28) is clear [3].

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